

Introducing the General Public to FITS Images

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Although fully-processed astronomical images in FITS format as well as the tools to work with these files are publicly available, neither most educators nor the general public are aware of the existence of these data or how to access them.

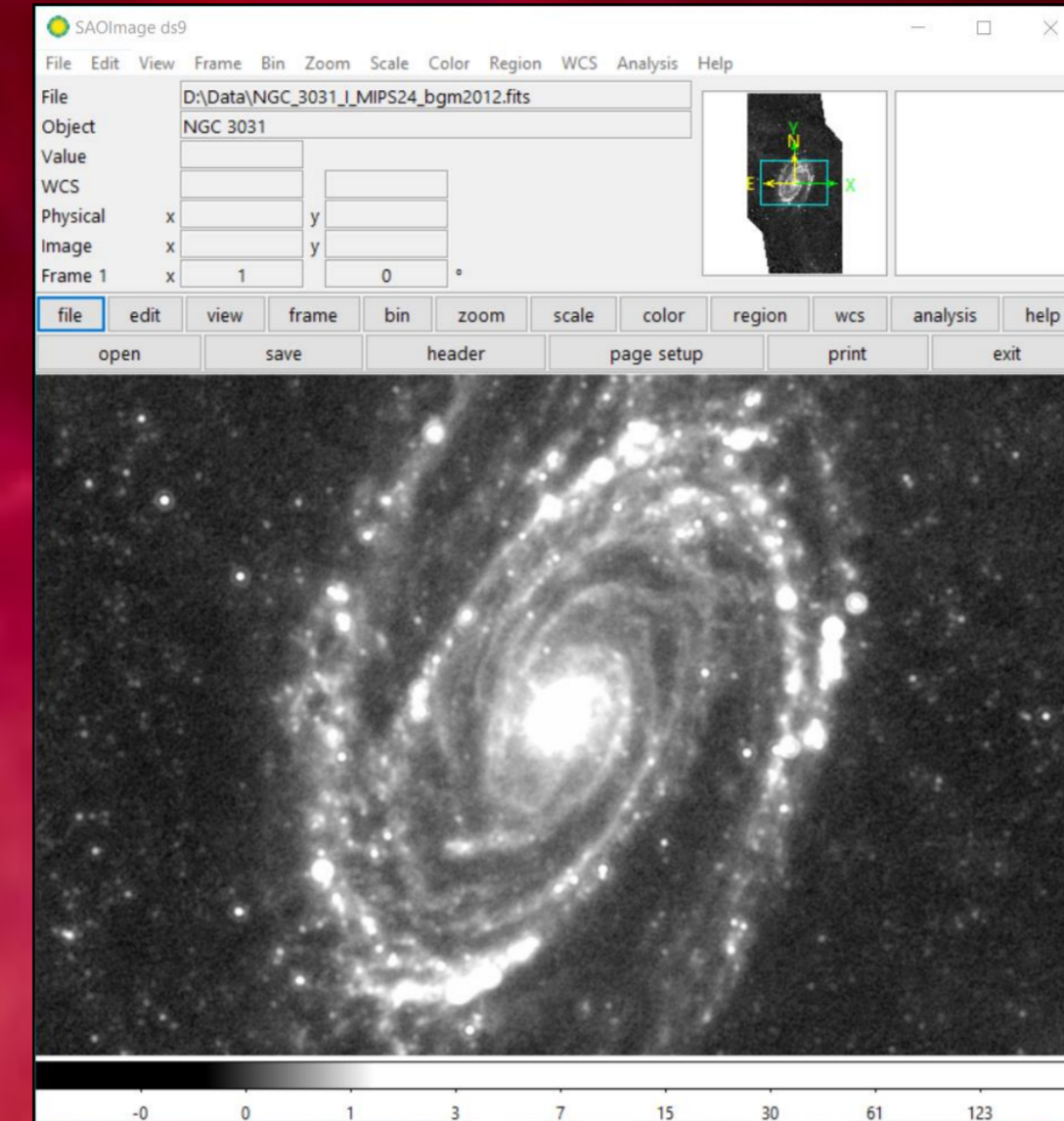
Through work with secondary school and university students, I have developed a series of activities that teach people how they can download astronomical FITS files and perform quantitative analyses with these data.

<http://www.jb.man.ac.uk/~gbendo/Sci/Teach/teachmaterial.html>

Visit the website above (scan the QC code for the website) to download the educational material on this poster.

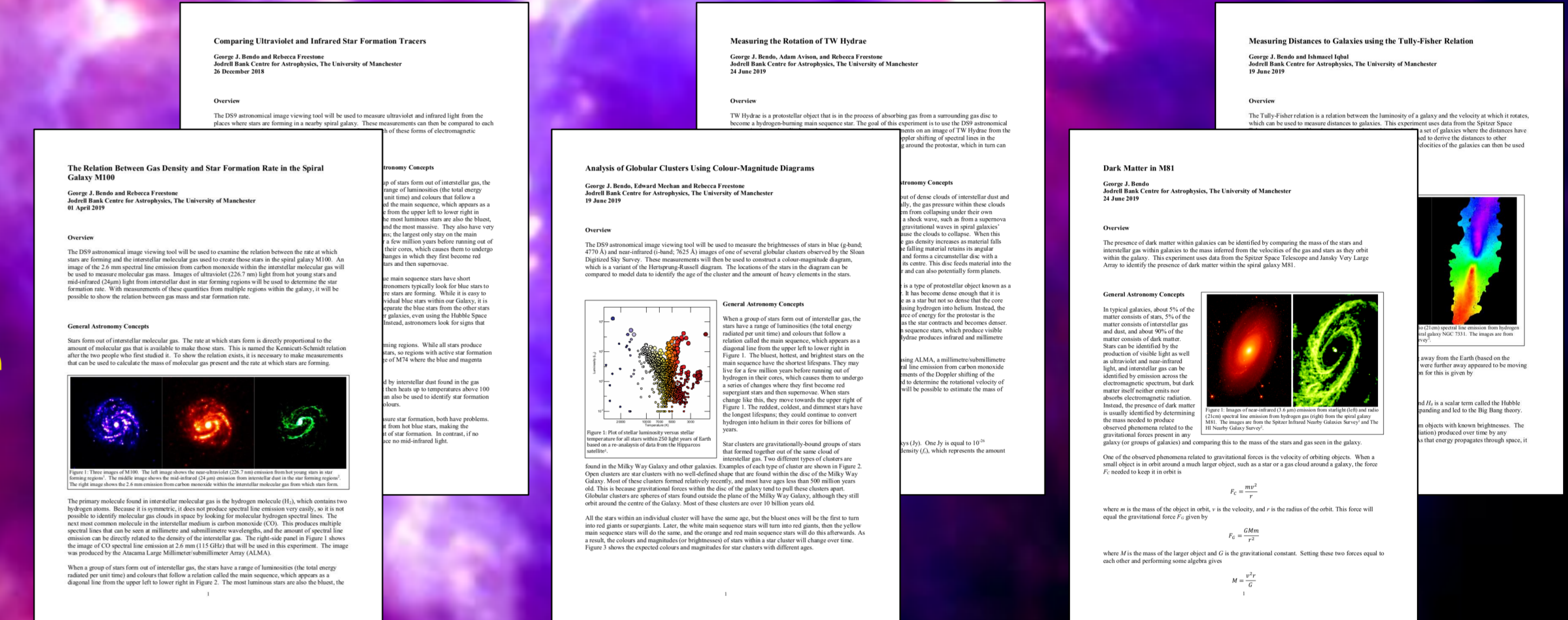


Introduction to Astronomy Images and the DS9 Image Viewer



This guide was developed to provide a straightforward introduction to working with FITS images and the DS9 image viewer. It can be used by secondary school students, university students, or the general public.

Lab Scripts for Secondary School and Undergraduate Teaching



These lab scripts were written (with the assistance of students) as classroom activities that could introduce students to using professional astronomy data. Scripts for secondary school and undergraduate students are currently available. The scripts have been tested with year 12 work experience students, and some are used for teaching at the University of Manchester.

Astronomers use right ascension and declination to identify locations in images.

Right ascension is the equivalent of longitude, but it is measured in hours instead of in degrees.

FITS files contain headers that may contain various information about the files.

FITS image viewers can be used to look at this information.

This information could include the following:

- Image creation information
- Coordinate information
- Data units
- Data processing history

Next, do the following:

- Move the cursor to the image window.
- Hold the right mouse button down.

Next, click on Scale.

After this, click on log.

The emission from the entire galaxy is now visible.

(With some images, clicking on Scale in the menu bar and then selecting 99.5%, Zscale, or Zmax sometimes produces better results.)

The introduction provides an overview of the FITS format and discusses basic concepts regarding astronomy images.

The quick start guide provides simple instructions on how to display images in DS9.

Adjust the scale and the brightness/contrast in the green channel so that it looks the way you want it to look.

Parts of the image with emission from both the red and green channels will look yellow.

If you double-click on the circle, an information box will appear with the coordinates and size of the circle.

You can also open this window by selecting the region in the image and then selecting "Get Information" from the region menu.

You can put the circle in a specific position or set it to a certain size if needed.

Websites with Professional FITS Image Archives

Professional astronomers use these websites to store and distribute their data, including FITS images. Many of these websites expect people to search for images of specific objects, a list of example objects is provided after this list of websites.

Some of these data may come in compressed formats such as tar, gzip, or bz2. These files can be uncompressed in Windows using 7Zip (http://www.7-zip.org) or 7-Zip (http://www.sevenzip.com). Mac and Linux computers have software built in to decompress these files.

It might also be necessary to add ".fits" to the ends of file to make them work in DS9.

- ESAO

Electromagnetic Spectrum

Band	Wavelengths	Emission Sources
Radio	>3 mm	• Supernovae • AGN
Millimetre	1 mm - 4 mm	• Atomic interstellar gas (hydrogen 21cm spectral line) • AGN
Submillimetre	250 µm - 1 mm	• Diatomic interstellar gas • Molecular interstellar gas (CO spectral line)
Far-infrared	50 µm - 500 µm	• Cold interstellar dust • Molecular interstellar gas (CO spectral line)
Mid-infrared	5 µm - 50 µm	• Hot interstellar dust • Large interstellar carbon molecules (PAHs)
Near-infrared	780 nm - 5 µm	• Old stars • Old stars (red wavelengths) • Young stars (blue wavelengths) • Warm ionized interstellar gas (H α line, H β line, other spectral lines)
Optical	380 - 780 nm	• Old stars (red wavelengths) • Young stars (blue wavelengths) • Warm ionized interstellar gas (H α line, H β line, other spectral lines)
Ultraviolet	10 nm - 380 nm	• Hot ionized interstellar gas • Many binary stars • AGN
X-ray	10 pm - 10 nm	• Hot ionized interstellar gas • AGN
Gamma-ray	> 50 pm	• Gamma ray bursts • AGN

The more detailed information on DS9 covers various menu options and discusses how to perform photometry and create multiband colour images.

The end of the guide lists links to websites with FITS images, suggested objects to look up on those websites, and information on the electromagnetic spectrum.

All lab scripts are based on using real science data in the form of FITS images acquired from public archives such as the NASA/IPAC Extragalactic Database and Sloan Digitized Sky Survey.

All analyses are based on using DS9, which is publicly available for free for all computer platforms.

The experiments cover topics such as the measuring stellar ages and metallicities for globular clusters and deriving the Kennicutt-Schmidt relation (the relation between gas mass and star formation rate).

The lab scripts not only provide introductions to these topics but also cover basic concepts such as the electromagnetic spectrum, Keplerian motion, and the expansion of the Universe.

